Einstein's velocity addition

("Zur Elektrodynamik bewegter Körper", Ann. Phys. 17 (1905), 891-921):

$$\begin{array}{ll} @ \ \mathrm{p. 906:} & U = \frac{\sqrt{\nu^2 + w^2 + 2\nu w \cos \alpha - \left(\frac{\nu w \sin \alpha}{c}\right)^2}}{1 + \frac{\nu w \cos \alpha}{c^2}} \\ \\ \text{dimensionless:} & \frac{U}{c} = \frac{\sqrt{\frac{\nu^2}{c^2} + \frac{w^2}{c^2} + 2\frac{\nu}{c} \frac{w}{c} \cos \alpha - \left(\frac{\nu}{c} \frac{w}{c} \sin \alpha\right)^2}}{1 + \frac{\nu}{c} \frac{w}{c} \cos \alpha} \\ \\ \text{which is:} & \beta_0 = \frac{\sqrt{\beta_1^2 + \beta_2^2 + 2\beta_1 \beta_2 \cos \alpha} - \beta_1^2 \beta_2^2 \sin^2 \alpha}{1 + \beta_1 \beta_2 \cos \alpha} \\ \\ \text{Note:} & \left\{\alpha, \beta_0, \beta_1\right\} \text{ observed in stationary frame,} \\ \\ \beta_2 \text{ in frame moving at } \beta_1 \text{ in stat. frame.} \\ \\ \text{Parallel:} & (\alpha = 0) \Rightarrow \quad \beta_0 = \frac{\sqrt{\beta_1^2 + \beta_2^2 + 2\beta_1 \beta_2}}{1 + \beta_1 \beta_2} = \frac{\beta_1 + \beta_2}{1 + \beta_1 \beta_2} \\ \\ \text{Perpendicular:} & \left(\alpha = \frac{\pi}{2}\right) \Rightarrow \quad \beta_0 = \sqrt{\beta_1^2 + \beta_2^2} - \beta_1^2 \beta_2^2 \end{array}$$

p.1/6

Lorentz factor:

$$\gamma = \frac{1}{\sqrt{1-\beta^2}} \therefore \gamma^2 (1-\beta^2) = 1 \therefore \gamma^2 - \gamma^2 \beta^2 = 1 \therefore \gamma^2 - 1 = \gamma^2 \beta^2$$
$$\therefore \beta^2 = \frac{\gamma^2 - 1}{\gamma^2} = 1 - \frac{1}{\gamma^2}$$

Perpendicular velocity addition:

$$\beta_0 = \sqrt{\beta_1^2 + \beta_2^2} - \beta_1^2 \beta_2^2$$

$$\boldsymbol{\gamma_0} = \frac{1}{\sqrt{1-\beta_0^2}} = \frac{1}{\sqrt{1-\beta_1^2 - \beta_2^2 + \beta_1^2 \beta_2^2}} = \frac{1}{\sqrt{1-\left(1-\frac{1}{\gamma_1^2}\right) - \left(1-\frac{1}{\gamma_2^2}\right) + \left(1-\frac{1}{\gamma_1^2}\right)\left(1-\frac{1}{\gamma_2^2}\right)}}$$

$$= \frac{1}{\sqrt{1-1+\frac{1}{\gamma_1^2} - 1+\frac{1}{\gamma_2^2} + \left(1-\frac{1}{\gamma_1^2}-\frac{1}{\gamma_2^2}+\frac{1}{\gamma_1^2}\cdot\frac{1}{\gamma_2^2}\right)}} = \frac{1}{\sqrt{\frac{1}{\gamma_1^2\gamma_2^2}}} = \boldsymbol{\gamma_1\gamma_2}$$

We have:

hence:

$$E_{\rm tot} = \gamma m c^2$$
$$\gamma = E_{\rm tot} / m c^2$$

which physically is:

dimensionless specific energy (i.e. energy per mass).

For perpendicular motions, the total specific energy is the **product** of the individual values!

$$\boldsymbol{\gamma_0} = \boldsymbol{\gamma_1} \boldsymbol{\gamma_2} = \frac{E_{\text{tot},1} E_{\text{tot},2}}{m_1 m_2 c^4} = \frac{(E_{\text{kin},1} + m_1 c^2)(E_{\text{kin},2} + m_2 c^2)}{m_1 m_2 c^4}$$

Parallel:

perpendicular:

$$B_{0} = \frac{\beta_{1} + \beta_{2}}{1 + \beta_{1} \beta_{2}}$$

$$B_{0} = \sqrt{\beta_{1}^{2} + \beta_{2}^{2}} - \beta_{1}^{2} \beta_{2}^{2} \quad \& \quad \gamma_{0} = \gamma_{1} \gamma_{2}$$

suffix 0:object's velocity w.r.t. stationary observer;suffix 2:object's velocity w.r.t. some moving ref. point P;suffix 1:P's velocity w.r.t. stationary observer

(consistent with the aforementioned, but 1 & 2 are of course interchangeable).

$$(\boldsymbol{\beta} \to \mathbf{0}) \Leftrightarrow (\boldsymbol{\gamma} \to \mathbf{1}) \qquad (\boldsymbol{\beta} \to \mathbf{1}) \Leftrightarrow (\boldsymbol{\gamma} \to \infty)$$

If $\{\beta_1, \gamma_1\} \rightarrow \{1, \infty\}$ The more an object's velocity w.r.t. P approaches c,then $\{\beta_0, \gamma_0\} \rightarrow \{1, \infty\}$ the more its velocity w.r.t. stat. obs. approaches c,totally indep. of $\{\beta_2, \gamma_2\}$ i.e. the more its velocity becomes absolute, namely c.

Relativity yields absoluteness...

Lorentz contraction of a moving object:

 $L \coloneqq$ length of moving object, measured in direction of motion;

 $L_{\text{seenByStationary}} = L_{\text{seenByComoving}} \sqrt{1 - \beta^2}$

If *you* are speeding, then from *your* perspective the street is moving, hence contracted by the same factor $\sqrt{1-\beta^2}$.

The faster you go, the shorter the street appears to you (cf. light experiencing zero time and distance).

Ever realised that street extends to "EDGE" of COSMOS?

<u>Would</u> you reach the speed of light, then you would — *at this very speed of light* — <u>instantly</u> smack against the "edge" of the universe! (Whatever the latter might be). Special Relativity presumes symmetry, both observers are equally equal. This premise implies the evident but implicit assumption

> that **YOUR** speed w.r.t. **ME**, as measured by **ME** equals **MY** speed w.r.t. **YOU**, as measured by **YOU**. $v_{me} = \Delta L_{me} / \Delta t_{me} = \Delta L_{you} / \Delta t_{you} = v_{you}$ (suffix means observed by, in hir own frame).

We usually take for granted

that	YOUR	speed w.r.t. <i>ME,</i>	as measured by YOURSELF
equals	YOUR	speed w.r.t. <i>ME,</i>	as measured by ME
AND	MY	speed w.r.t. YOU ,	as measured by MYSELF
equals	MY	speed w.r.t. YOU ,	as measured by YOU,

which inevitably renders *Lorentz contraction*:

due to time dilation, moving observer perceives shorter distance. Would we however reject this presumption we merely took for granted, then both perceive the same length of the street, but the traveller needs less time as perceived by hirself, so (s)he perceives hir own velocity as way larger than how the stat. obs. perceives it. From hir own perspective, (s)he could go way faster than the speed of light! It is called celerity and it equals: $\Delta L_{stat} / \Delta t_{mov} = \gamma \beta c$.